## WikipediA

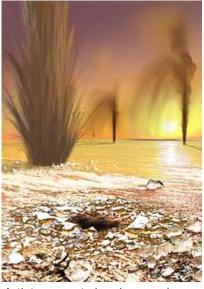
# **Geysers on Mars**

**Martian geysers** (or  $CO_2$  **jets**) are putative sites of small gas and dust eruptions that occur in the <u>south polar region</u> of <u>Mars</u> during the spring thaw. "Dark dune spots" and "spiders" – or araneiforms [1] – are the two most visible types of features ascribed to these eruptions.

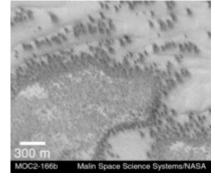
Martian geysers are distinct from geysers on Earth, which are typically associated with hydrothermal activity. These are unlike any terrestrial geological phenomenon. The reflectance (albedo), shapes and unusual spider appearance of these features have stimulated a variety of hypotheses about their origin, ranging from differences in frosting reflectance, to explanations involving biological processes. However, all current geophysical models assume some sort of jet or geyser-like activity on Mars. [2][3][4][5][6][7][8][9][10] Their characteristics, and the process of their formation, are still a matter of debate.

These features are unique to the south polar region of Mars in an area informally called the 'cryptic region', at latitudes  $60^{\circ}$  to  $80^{\circ}$  south and longitudes  $150^{\circ}$ W to  $310^{\circ}$ W; [11][12][13] this 1 meter deep carbon dioxide (CO<sub>2</sub>) ice transition area—between the scarps of the thick polar ice layer and the permafrost—is where clusters of the apparent geyser systems are located.

The seasonal frosting and defrosting of carbon dioxide ice results in the appearance of a number of features, such dark dune spots with spider-like <u>rilles</u> or channels below the ice, [3] where spider-like radial channels are carved between the ground and the carbon dioxide ice, giving it an appearance of spider webs, then, pressure accumulating in their interior ejects gas and dark basaltic sand or dust, which is deposited on the ice surface and thus, forming dark dune spots. [2][3][4][5][6][7][8] This process is rapid, observed happening in the space of a few days, weeks or months, a growth rate rather unusual in geology – especially for Mars. [14] However, it would seem that multiple years would be required to carve the larger spider-like channels. [2] There is no direct data on these features other than images taken in the <u>visible</u> and infrared spectra.



Artist concept showing sandladen jets erupting from Martian geysers. (Published by NASA; artist: Ron Miller.)



Dark dune spots

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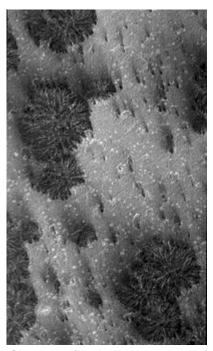
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## History

The geological features informally called **dark dune spots** and **spiders** were separately discovered on images acquired by the MOC camera on board the Mars Global Surveyor during 1998–1999. [15][16] At first it was generally thought they were unrelated features because of their appearance, so from 1998 through 2000 they were reported separately on different research publications ([16][17]] and [18] - respectively). "Jet" or "geyser" models were proposed and refined from 2000 onwards. [4][5]

The name 'spiders' was coined by Malin Space Science Systems personnel, the developers of the camera. One of the first and most interesting spider photos was found by Greg Orme in October 2000. The unusual shape and appearance of these 'spider webs' and spots caused a lot of speculation about their origin. The first years' surveillance showed that during the following Martian years, 70% of the spots appear at exactly the same place, and a preliminary statistical study obtained between September 1999 and March 2005, indicated that dark dune spots and spiders are related phenomena as functions of the cycle of carbon dioxide ( $\rm CO_2$ ) condensing as " $\rm \underline{dry}$  ice" and sublimating.  $\rm \underline{[20]}$ 

It was also initially suggested that the dark spots were simply warm patches of bare ground, but thermal imaging during 2006 revealed that these structures were as cold as the ice that covers the area, [9][20]



Close up of dark dune spots obtained by the Mars Global Surveyor and discovered in 2000 by Greg Orme.

indicating they were a thin layer of dark material lying on top of the ice and kept chilled by it. [9] However, soon after their first detection, they were discovered to be negative topographical features – i.e. radial troughs or channels of what today are thought to be geyser-like vent systems. [2][3][4][5][6][7][8]

## **Morphology**

The geysers' two most prominent features (dark dune spots and spider channels) appear at the beginning of the Martian spring on dune fields covered with carbon dioxide (CO<sub>2</sub> or 'dry ice'), mainly at the ridges and slopes of the dunes; by the beginning of winter, they disappear. Dark spots' shape is generally round, on the slopes it is usually elongated, sometimes with streams—possibly of water—that accumulate in

pools at the bottom of the dunes. [21][22] Dark dune spots are typically 15 to 46 metres (50 to 150 feet) wide and spaced several hundred feet apart. [9] The size of spots varies, and some are as small as 20 m across, [16][23]—however, the smaller size seen is limited by imaging resolution—and can grow and coalesce into formations several kilometres wide.

Spider features, when viewed individually, form a round lobed structure reminiscent of a spider web radiating outward in lobes from a central point. [24] Its radial patterns represent shallow channels or ducts in the ice formed by the flow of the sublimation gas toward the vents. [3][4] The entire spider channel network is typically 160–300 m across, although there are large variations. [2]

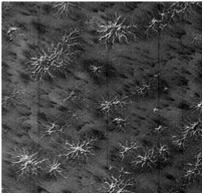
Each geyser's characteristic form appears to depend on a combination of such factors as local fluid or gas composition and pressure, ice thickness, underlying gravel type, local climate and meteorological conditions. The geysers' boundary does not seem to correlate with any other properties of the surface such as elevation, geological structure, slope, chemical composition or thermal properties. The geyser-like system produce low-albedo spots, fans and blotches, with small radial spider-like channel networks most often associated with their location. At first, the spots seem to be grey, but later their centres darken because they gradually get covered with dark ejecta, thought to be mainly basaltic sand. Not all dark spots observed in early spring are associated with spider landforms, however, a preponderance of dark spots and streaks on the cryptic terrain are associated with the appearance of spiders later in the season.

Time-lapsed imagery performed by NASA confirms the apparent ejection of dark material following the radial growth of spider channels in the ice. [9] Time-lapsed imaging of a single area of interest also shows that small dark spots generally indicate the position of spider features not yet visible; it also shows that spots expand significantly, including dark fans emanating from some of the spots, which increase in prominence and develop clear directionality indicative of wind action [2]

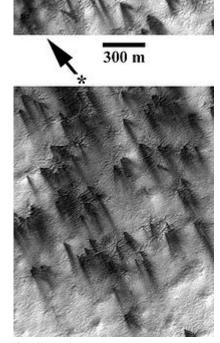
Some branching ravines modify, some destroy and others create crust in a dynamic near-surface process that extensively reworks the terrain creating and destroying surface layers. Thus, Mars seems to have a dynamic process of recycling of its near surface crust of carbon dioxide. Growth process is rapid, happening in the space of a few days, weeks or months, a growth rate rather unusual in geology – especially for Mars. [14] A number of geophysical models have been investigated to explain the various colors and shapes' development of these geysers on the southern polar ice cap of Mars.



Dark dune spots. High resolution color image by the <u>HiRISE</u> camera



'Spider' features shown in relationship to dark dune spots.



Dark sediment spots apparently emanating from 'spider' formations.

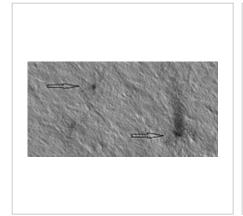


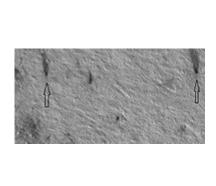




Wide view of plumes, as seen by HiRISE under HiWish under HiWish program Arrow HiRISE under HiWish program program Many of the plumes shows a double plume. This show spiders when enlarged.

Plumes, as seen by HiRISE Long plume, as seen by may have been because of shifting winds.







under HiWish program

by HiRISE under HiWish by program

Spiders, as seen by HiRISE Plumes and spiders, as seen Plumes and spiders, as seen **HIRISE** under HiWish program







Plumes and spiders, as seen Wide view of plumes and Plumes and spiders, as seen under HiWish by HiRISE program

spiders, as seen by HiRISE by under HiWish program

HiRISE under program



Spiderlike terrain on Mars

## Geyser mechanism models

The strength of the eruptions is estimated to range from simple upsurges to high-pressure eruptions at speeds of 160 kilometres per hour (99 mph) or more, [4][25] carrying dark basaltic sand and dust plumes high aloft. [9] The current proposed models dealing with the possible forces powering the geyser-like system are discussed next.

### Atmospheric pressure

The surface atmospheric pressure on Mars varies annually around: 6.7–8.8 mbar and 7.5–9.7 mbar; daily around 6.4-6.8 mbar. Because of the pressure changes subsurface gases expand and contract periodically, causing a downward gas flow during increase of and expulsion during decrease of atmospheric pressure. [7] This cycle was first quantified with measurements of the surface pressure, which varies annually with amplitude of 25%.[2]

### Clathrate hydrate model

This model proposes downward gas flow during increase of and upward flow during decrease of atmospheric pressure. In the defrosting process, ices (clathrate) may partly migrate into the soil and partly may evaporate. These locations can be in connection with the formation of dark dune spots and the arms of spiders as gas travel paths.

### **Dry venting**

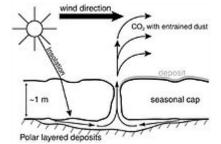
Some teams propose dry venting of carbon dioxide (CO<sub>2</sub>) gas and sand, occurring between the ice and the underlying bedrock. It is known that a CO2 ice slab is virtually transparent to solar radiation where 72% of solar energy incident at 60 degrees off vertical will reach the bottom of a 1 m thick layer. [4][27] In addition, separate teams from Taiwan and France measured the ice thickness in several target areas, and discovered that the greatest thickness of the CO2 frost layer in the geysers' area is about 0.76-0.78 m, supporting the geophysical model of dry venting powered by sunlight. [8][28][29] As the southern spring CO<sub>2</sub> ice receives enough solar energy, it starts sublimation of the CO<sub>2</sub> ice from the bottom. [2] This vapor accumulates under the slab rapidly increasing pressure and erupting. [6][9][14][30][31] High-pressure gas flows through at speeds of 160 kilometres per hour (99 mph) or more; [4][25] under the slab, the gas erodes ground as it rushes toward the vents, snatching up loose particles of sand and carving the spidery network of grooves.[8] The dark material falls back to the surface and may be taken up slope by wind, creating dark wind streak patterns on the ice cap. [20][25] This model is consistent observations. [25][32] The location, size and direction of these fans are useful to quantifying seasonal winds and sublimation activity. [26]

It is clear that sublimation of the base of the seasonal ice cap is more than capable of generating a substantial overpressure, [2] which is four orders of magnitude higher than the ice overburden pressure and five orders of magnitude higher than atmospheric pressure as discussed above.<sup>[2]</sup>

The observation that a few dark spots form before sunrise, with significant spot formation occurring immediately following sunrise, supports the notion that the system is powered by solar energy. [33] Eventually the ice is completely removed and the dark granular material is back on the surface; [33] the cycle repeats many times. [20][34][35]



A large 'spider' feature apparently emanating sediment to give rise to dark dune spots. Image size: 1 km (0.62 mi) across.



According to Sylvain Piqueux, sun light causes sublimation from the bottom, leading to a buildup of pressurized CO<sub>2</sub> gas which eventually bursts out, entraining dust and leading to dark fan-shaped deposits with clear directionality indicative of wind action. [26]

Laboratory experiments performed in 2016 were able to trigger dust eruptions from a layer of dust inside a  $\rm CO_2$  ice slab under Martian atmospheric conditions, lending support to the  $\rm CO_2$  jet and fan production model. [26]

#### Water-driven erosion

Data obtained by the Mars Express satellite, made it possible in 2004 to confirm that the southern polar cap has an average of 3 kilometres (1.9 mi) thick slab of  $CO_2$  ice<sup>[36]</sup> with varying contents of frozen water, depending on its latitude: the bright polar cap itself, is a mixture of 85%  $CO_2$  ice and 15% water ice. [37] The second part comprises steep slopes known as 'scarps', made almost entirely of water ice, that fall away from the polar cap to the surrounding plains. [37] This transition area between the scarps and the permafrost is the 'cryptic region', where clusters of geysers are located.

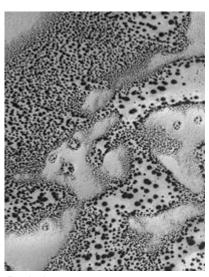
This model explores the possibility of active water-driven erosive structures, where soil and water derived from the shallow sub-surface layer is expelled up by  $CO_2$  gas through fissures eroding joints to create spider-like radiating tributaries capped with mud-like material and/or ice. [14][38][39][40]

#### **Geothermal**

A European team proposes that the features could be a sign that non-solar energy source is responsible of the jets, subsurface heat wave for instance. This model is difficult to reconcile with the evidence collected in the form of thermal emission (infrared) imaging, which shows that the fans, spots and blotches are produced by expulsion of cold fluids or cold gases. [31][42]

## Carbon dioxide and water cycling

Michael C. Malin, a planetary scientist and designer of the cameras used by the Mars Global Surveyor that obtained the earliest images of the CO<sub>2</sub> geyser phenomenon, is studying the images acquired of specific areas and he tracks their changes over a period of a few years. In 2000, he modelled the fans and spots' dynamics as a complex process of carbon dioxide (CO<sub>2</sub>) and water sublimation and reprecipitation. The typical pattern of defrosting proceeds from the initiation of small, dark spots typically located at the margins of dunes; these spots individually enlarge and eventually all coalesce. [34] The pattern the enlargement follows is distinct and characteristic: a dark nuclear spot enlarges slowly, often with a bright outer zone or 'halo'. As these are progressive, centripetal phenomena, each location of the light zone is overtaken by an expanding dark zone. Although initially developed along dune margins, spot formation quickly spreads onto and between dunes. As spring progresses, fan-shaped tails ('spiders') develop from the central spot. Defrosting occurs as the low albedo polar



Dark dune spots

sand heats beneath an optically thin layer of frost, causing the frost to evaporate. This is the dark nucleus of the spots seen on dunes. As the vapor moves laterally, it encounters cold air and precipitates, forming the bright halo. This precipitated frost is again vaporized as the uncovered zone of sand expands; the cycle repeats many times. [20][34][35]

## **European Space Agency**

While the <u>European Space Agency</u> (ESA) has not yet formulated a theory or model, they have stated that the process of frost sublimation is not compatible with a few important features observed in the images, and that the location and shape of the spots is at odds with a physical explanation, specifically, because the channels appear to radiate downhill as much as they radiate uphill, defying gravity. [43]

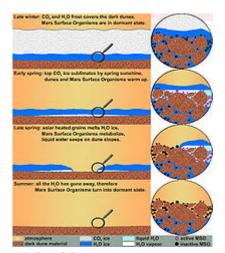
## Hypothetical biological origin

A team of Hungarian scientists propose that the dark dune spots and channels may be colonies of photosynthetic Martian microorganisms, which over-winter beneath the ice cap, and as the sunlight returns to the pole during early spring, light penetrates the ice, the microorganisms photosynthesise and heat their immediate surroundings. A pocket of liquid water, which would normally evaporate instantly in the thin Martian atmosphere, is trapped around them by the overlying ice. As this ice layer thins, the microorganisms show through grey. When it has completely melted, they rapidly desiccate and turn black surrounded by a grey aureole. [22][44][45][46] The Hungarian scientists think that even a complex sublimation process is insufficient to explain the formation and evolution of the dark dune spots in space and time. [23][47] Since their discovery, fiction writer Arthur C. Clarke promoted these formations as deserving of study from an astrobiological perspective. [19]

A multinational European team suggests that if liquid water is present in the spiders' channels during their annual defrost cycle, the structures might provide a niche where certain microscopic life forms could have retreated and adapted while sheltered from <u>UV</u> solar radiation. [3] British and German teams also consider the possibility that organic



Dark dune spots.



DDS-MSO hypothesis.

matter, microbes, or even simple plants might co-exist with these inorganic formations, especially if the mechanism includes liquid water and a geothermal energy source. [14][48] However, they also remark that the majority of geological structures may be accounted for without invoking any organic "life on Mars" hypothesis. [14] (See also: Life on Mars.)

## **Lander mission**

There is no direct data on these features other than images taken in the <u>visible</u> and infrared spectra, and development of the <u>Mars Geyser Hopper lander</u> is under consideration to study the geyser-like systems. [49][50] It has not yet been formally proposed nor funded.

## See also

- Arachnoid (astrogeology) Large structure of unknown origin on Venus
- Chaos terrain Distinctive area of broken or jumbled terrain
- Geology of Mars Scientific study of the surface, crust, and interior of the planet Mars
- Planetary geology The geology of astronomical objects apparently in orbit around stellar objects
- Rille Fissure, especially on the Moon
- Swiss cheese features Enigmatic surface features on Mars southern ice cap

### References

1. Portyankina, Ganna (2014). "Araneiform". *Encyclopedia of Planetary Landforms*. p. 1. doi:10.1007/978-1-4614-9213-9\_540-1 (https://doi.org/10.1007%2F978-1-4614-9213-9\_540-1). ISBN 978-1-4614-9213-9. Missing or empty |title= (help)

- Piqueux, Sylvain; Shane Byrne; Mark I. Richardson (8 August 2003). "Sublimation of Mars's southern seasonal CO<sub>2</sub> ice cap formation of spiders" (http://www.lpl.arizona.edu/~shane/publication s/piqueux\_etal\_jgr\_2003.pdf) (PDF). Journal of Geophysical Research. 180 (E8): 5084. Bibcode:2003JGRE..108.5084P (https://ui.adsabs.harvard.edu/abs/2003JGRE..108.5084P). doi:10.1029/2002JE002007 (https://doi.org/10.1029%2F2002JE002007).
- 3. Manrubia, S. C.; O. Prieto Ballesteros; C. González Kessler; D. Fernández Remolar; C. Córdoba-Jabonero; F. Selsis; S. Bérczi; T. Gánti; A. Horváth; A. Sik; E. Szathmáry (2004). "Comparative Analysis of Geological Features and Seasonal Processes in Inca City and PittyUSA Patera Regions on Mars" (https://web.archive.org/web/20110721105008/http://www.colbud.hu/esa/publications/29ProcCAB-3ESASP-545pp77-80.pdf) (PDF). European Space Agency Publications (ESA SP): 545. Archived from the original (http://www.colbud.hu/esa/publications/29ProcCAB-3ESASP-545pp77-80.pdf) (PDF) on 21 July 2011.
- Kieffer, H. H. (2000). "Mars Polar Science 2000 Annual Punctuated CO<sub>2</sub> Slab-ice and Jets on Mars" (http://www.lpi.usra.edu/meetings/polar2000/pdf/4095.pdf) (PDF). Retrieved 6 September 2009.
- Kieffer, Hugh H. (2003). "Third Mars Polar Science Conference (2003)- Behavior of Solid CO" (http://www.lpi.usra.edu/meetings/polar2003/pdf/8083.pdf) (PDF). Retrieved 6 September 2009.
- 6. Portyankina, G., ed. (2006). "Fourth Mars Polar Science Conference Simulations of Geyser-Type Eruptions in Cryptic Region of Martian South" (http://www.lpi.usra.edu/meetings/polar2006/pdf/8040.pdf) (PDF). Retrieved 11 August 2009.
- Sz. Bérczi; et al., eds. (2004). "Lunar and Planetary Science XXXV (2004) Stratigraphy of Special Layers – Transient Ones on Permeable Ones: Examples" (http://www.lpi.usra.edu/meetings/lpsc200 4/pdf/1317.pdf) (PDF). Retrieved 12 August 2009.
- 8. Kieffer, Hugh H.; Philip R. Christensen; Timothy N. Titus (30 May 2006). "CO<sub>2</sub> jets formed by sublimation beneath translucent slab ice in Mars' seasonal south polar ice cap". *Nature*. 442 (7104): 793–6. Bibcode: 2006Natur.442..793K (https://ui.adsabs.harvard.edu/abs/2006Natur.442..793K). doi:10.1038/nature04945 (https://doi.org/10.1038%2Fnature04945). PMID 16915284 (https://pubmed.ncbi.nlm.nih.gov/16915284). S2CID 4418194 (https://api.semanticscholar.org/CorpusID:4418194).
- "NASA Findings Suggest Jets Bursting From Martian Ice Cap" (http://www.jpl.nasa.gov/news/news.cf m?release=2006-100). Jet Propulsion Laboratory. NASA. 16 August 2006. Retrieved 11 August 2009.
- C.J. Hansen, N. Thomas, G. Portyankina, A. McEwen, T. Becker, S. Byrne, K. Herkenhoff, H. Kieffer, M. Mellon (2010). "HiRISE observations of gas sublimation-driven activity in Mars' southern polar regions: I. Erosion of the surface" (http://www.lpl.arizona.edu/~shane/publications/hansen\_etal\_icaru s\_2010.pdf) (PDF). *Icarus*. 205 (1): 283–295. Bibcode:2010Icar..205..283H (https://ui.adsabs.harvard.edu/abs/2010Icar..205..283H). doi:10.1016/j.icarus.2009.07.021 (https://doi.org/10.1016%2Fj.icarus.2009.07.021).
- 11. Titus T. N. et al. (2003) Third Mars Polar Science Conference, Abstract #8081.
- 12. Kieffer, H. H. (2001) Second International Conf. On Mars Polar Sci. and Exploration, no. 1057.
- 13. Kieffer, H. H. (2003), Sixth International Conference on Mars, no. 3158.
- 14. Ness, Peter K.; Greg M. Orme (2002). "Spider-Ravine Models and Plant-like Features on Mars Possible Geophysical and Biogeophysical Modes of Origin" (https://web.archive.org/web/201202201 30910/http://spsr.utsi.edu/articles/ness.pdf) (PDF). Journal of the British Interplanetary Society (JBIS). 55: 85–108. Archived from the original (http://spsr.utsi.edu/articles/ness.pdf) (PDF) on 20 February 2012. Retrieved 3 September 2009.
- Albee, A. L.; F. D. Palluconi; R. E. Arvidson (1998). "Mars Global Surveyor Mission: Overview and Status" (http://www.sciencemag.org/cgi/content/full/279/5357/1671). Science. 279 (5357): 1671–5. Bibcode: 1998Sci...279.1671A (https://ui.adsabs.harvard.edu/abs/1998Sci...279.1671A). doi:10.1126/science.279.5357.1671 (https://doi.org/10.1126%2Fscience.279.5357.1671). PMID 9497277 (https://pubmed.ncbi.nlm.nih.gov/9497277).

- 16. Malin, Michael C.; et al. (13 March 1998). "Early Views of the Martian Surface from the Mars Orbiter Camera of Mars Global Surveyor" (http://www.sciencemag.org/cgi/content/full/279/5357/1681). Science. 279 (5357): 1681–5. Bibcode: 1998Sci...279.1681M (https://ui.adsabs.harvard.edu/abs/1998 Sci...279.1681M). doi:10.1126/science.279.5357.1681 (https://doi.org/10.1126%2Fscience.279.5357.1681). PMID 9497280 (https://pubmed.ncbi.nlm.nih.gov/9497280).
- 17. Vasavada, A.; K. E. Herkenhoff (1999). "Surface Properties of Mars' Polar Layered Deposits and Polar Landing Sites" (http://mars.jpl.nasa.gov/mgs/sci/fifthconf99/6086.pdf) (PDF). NASA. Retrieved 21 August 2008.
- Lovett, R. A. (15 September 2000). "'Spiders' Channel Mars Polar Ice Cap". Science. 289 (5486): 1853a–4a. doi:10.1126/science.289.5486.1853a (https://doi.org/10.1126%2Fscience.289.5486.1853a). PMID 17839924 (https://pubmed.ncbi.nlm.nih.gov/17839924). S2CID 39054349 (https://api.semanticscholar.org/CorpusID:39054349).
- Orme, Greg M.; Peter K. Ness (9 June 2003). "Marsbugs" (https://web.archive.org/web/2009032713 5109/http://www.lyon.edu/projects/marsbugs/2003/20030609.pdf) (PDF). The Electronic Astrobiology Newsletter. 10 (23): 5. Archived from the original (http://www.lyon.edu/projects/marsbugs/2003/2003 0609.pdf) (PDF) on 27 March 2009. Retrieved 6 September 2009.
- J. J. Jian; W. H. Ip, eds. (2006). "Lunar and Planetary Science XXXVII (2006) Observation of the Martian Cryptic Region from Mars Orbiter Camera" (http://www.lpi.usra.edu/meetings/lpsc2006/pdf/1 313.pdf) (PDF). Retrieved 4 September 2009.
- Horváth, A.; Kereszturi, Á.; Bérczi, Sz.; et al. (2005). "Annual change of Martian DDS-seepages" (htt p://www.lpi.usra.edu/meetings/lpsc2005/pdf/1128.pdf) (PDF). Lunar and Planetary Science XXXVI: 1128. Bibcode: 2005LPI....36.1128H (https://ui.adsabs.harvard.edu/abs/2005LPI....36.1128H). Retrieved 24 November 2008.
- 22. Gánti, Tibor; András Horváth; Szaniszló Bérczi; Albert Gesztesi; Eörs Szathmáry (12–16 March 2001). "Probable Evidences of Recent Biological Activity on Mars: Appearance and Growing of Dark Dune Spots in the South Polar Region" (http://www.lpi.usra.edu/meetings/lpsc2001/pdf/1543.pdf) (PDF). 32nd Annual Lunar and Planetary Science Conference, Houston, Texas, Abstract No.1543: 1543. Bibcode:2001LPI....32.1543H (https://ui.adsabs.harvard.edu/abs/2001LPI....32.1543H). Retrieved 20 November 2008.
- A. Horváth; T. Gánti; Sz. Bérczi; A. Gesztesi; E. Szathmáry, eds. (2002). "Lunar and Planetary Science XXXIII - Morphological Analysis of the Dark Dune Spots on Mars: New Aspects in Biological Interpretation" (http://www.lpi.usra.edu/meetings/lpsc2002/pdf/1108.pdf) (PDF). Retrieved 24 November 2008.
- 24. "Spiders on Earth and Mars" (https://web.archive.org/web/20091013121837/http://www.aig.asn.au/pd f/AIGnews\_Aug06.pdf) (PDF). Australian Institute of Geoscientists. August 2006. p. 21. Archived from the original (http://www.aig.asn.au/pdf/AIGnews\_Aug06.pdf) (PDF) on 13 October 2009. Retrieved 11 August 2009.
- 25. Edgett, Kenneth S. (13 June 2002). "Low-albedo surfaces and eolian sediment: Mars Orbiter Camera" (https://web.archive.org/web/20040505233200/http://www.msss.com/biographies/edgett/reprints/2001JE001587.pdf) (PDF). Journal of Geophysical Research. 107 (E6): 5038. Bibcode:2002JGRE..107.5038E (https://ui.adsabs.harvard.edu/abs/2002JGRE..107.5038E). doi:10.1029/2001JE001587 (https://doi.org/10.1029%2F2001JE001587). hdl:2060/20010069272 (https://hdl.handle.net/2060%2F20010069272). Archived from the original (http://www.msss.com/biographies/edgett/reprints/2001JE001587.pdf) (PDF) on 5 May 2004.
- 26. Aye, K.-Michael; Schwamb, Megan E.; Portyankina, Ganna; et al. (2018). "Planet Four: Probing springtime winds on Mars by mapping the southern polar CO2 jet deposits" (https://zenodo.org/record/1475479). Icarus. 319: 558–598. doi:10.1016/j.icarus.2018.08.018 (https://doi.org/10.1016%2Fj.icarus.2018.08.018). ISSN 0019-1035 (https://www.worldcat.org/issn/0019-1035). S2CID 119103435 (https://api.semanticscholar.org/CorpusID:119103435).

- 27. Mangold, N (2011). "Ice sublimation as a geomorphic process: A planetary perspective". Geomorphology. 126 (1–2): 1–17. Bibcode:2011Geomo.126....1M (https://ui.adsabs.harvard.edu/abs/2011Geomo.126....1M). doi:10.1016/j.geomorph.2010.11.009 (https://doi.org/10.1016%2Fj.geomorph.2010.11.009).
- 28. Jian, Jeng-Jong; Ip, Wing-Huen (5 January 2009). "Seasonal patterns of condensation and sublimation cycles in the cryptic and non-cryptic regions of the South Pole". *Advances in Space Research*. **43** (1): 138–142. Bibcode:2009AdSpR..43..138J (https://ui.adsabs.harvard.edu/abs/2009AdSpR..43..138J). doi:10.1016/j.asr.2008.05.002 (https://doi.org/10.1016%2Fj.asr.2008.05.002).
- 29. Pilorget, C. (May 2011). "Dark spots and cold jets in the polar regions of Mars: New clues from a thermal model of surface CO2 ice" (https://hal.archives-ouvertes.fr/hal-00743841/file/PEER\_stage2\_10.1016%252Fj.icarus.2011.01.031.pdf) (PDF). *Icarus*. 213 (1): 131. Bibcode:2011Icar..213..131P (https://ui.adsabs.harvard.edu/abs/2011Icar..213..131P). doi:10.1016/j.icarus.2011.01.031 (https://doi.org/10.1016%2Fj.icarus.2011.01.031).
- 30. Hoffman, Nick (August 2002). "Active Polar Gullies on Mars and the Role of Carbon Dioxide". *Astrobiology*. **2** (3): 313–323. Bibcode:2002AsBio...2..313H (https://ui.adsabs.harvard.edu/abs/2002 AsBio...2..313H). doi:10.1089/153110702762027899 (https://doi.org/10.1089%2F153110702762027899). PMID 12530241 (https://pubmed.ncbi.nlm.nih.gov/12530241).
- 31. Titus, T. N.; Kieffer, H H; Langevin, Y; Murchie, S; Seelos, F; Vincendon, M; TEAM, C. (2007). "Bright Fans in Mars Cryptic Region Caused by Adiabatic Cooling of CO<sub>2</sub> Gas Jets" (https://www.researchg ate.net/publication/253157564). Eos, Transactions, American Geophysical Union. 88 (52 (Fall Meet. Suppl.)): P24A–05. Bibcode: 2007AGUFM.P24A..05T (https://ui.adsabs.harvard.edu/abs/2007AGUFM.P24A..05T).
- 32. Titus, T. N.; H. H. Kieffer; J. J. Plaut; P. R. Christensen; A. B. Ivanov; the THEMIS Science Team. (2003). "Third Mars Polar Science Conference (2003) South Polar Cryptic Region Revisited: THEMIS Observations" (http://www.lpi.usra.edu/meetings/polar2003/pdf/8081.pdf) (PDF). Retrieved 4 September 2009.
- 33. Kieffer, H H, Titus, T N, Christensen, P R (2005). "Infrared and Visible Observations of South Polar Spots and Fans" (https://web.archive.org/web/20090315074210/http://www.agu.org/cgi-bin/SFgate/SFgate). Eos, Transactions, American Geophysical Union. 86 (52 (Fall Meet. Suppl.)): P23C–04. Bibcode: 2005AGUFM.P23C..04C (https://ui.adsabs.harvard.edu/abs/2005AGUFM.P23C..04C). Archived from "P23C-04" the original (http://www.agu.org/cgi-bin/SFgate/SFgate?&listenv=table&multiple=1&range=1&directget=1&application=fm05&database=%2Fdata%2Fepubs%2Fwais%2Findexes%2Ffm05%2Ffm05&maxhits=200&=) on 15 March 2009. Retrieved 8 September 2009.
- 34. Malin, Michael C.; K. S. Edgett (2000). "Frosting and Defrosting of Martian Polar Dunes". <u>Lunar and Planetary Science XXXI</u> (http://www.lpi.usra.edu/meetings/lpsc2000/pdf/1056.pdf) (PDF). Malin Space Science Systems. Retrieved 3 September 2009.
- 35. Jeng-Jong Jian; Wing-Huen Ipa; Shin-Reu Sheu (2009). "Spatial distributions and seasonal variations of features related to a venting process at high southern latitudes observed by the MOC camera". *Planetary and Space Science*. **57** (7): 797–803. <u>Bibcode</u>:2009P&SS...57..797J (https://ui.a dsabs.harvard.edu/abs/2009P&SS...57..797J). doi:10.1016/j.pss.2009.02.014 (https://doi.org/10.1016%2Fj.pss.2009.02.014).
- 36. "Mars' South Pole Ice Deep and Wide" (http://jpl.nasa.gov/news/news.cfm?release=2007-030). *Jet Propulsion Laboratory*. NASA. 15 March 2007. Retrieved 11 September 2009.
- 37. "Water at Martian south pole" (http://www.esa.int/SPECIALS/Mars\_Express/SEMYKEX5WRD\_0.htm l). European Space Agency (ESA). 17 March 2004. Retrieved 11 September 2009.
- 38. Prieto-Ballesteros, Olga; Fernández-Remolar, DC; Rodríguez-Manfredi, JA; Selsis, F; Manrubia, SC (August 2006). "Spiders: Water-Driven Erosive Structures in the Southern Hemisphere of Mars". *Astrobiology*. **6** (4): 651–667. Bibcode:2006AsBio...6..651P (https://ui.adsabs.harvard.edu/abs/2006AsBio...6..651P). doi:10.1089/ast.2006.6.651 (https://doi.org/10.1089%2Fast.2006.6.651). PMID 16916289 (https://pubmed.ncbi.nlm.nih.gov/16916289).

- 39. Prieto-Ballesteros, Olga (2005). "Martian Spiders as feasible water-driven erosive structures" (https://web.archive.org/web/20090706072920/http://nai.nasa.gov/nai2005/abstracts/653%20-%20NAI2005-spiders.doc.pdf) (PDF). Centro de Astrobiología-INTA-CSIC. Archived from the original (http://nai.nasa.gov/nai2005/abstracts/653%20-%20NAI2005-spiders.doc.pdf) (PDF) on 6 July 2009. Retrieved 11 August 2009.
- Horváth, András; Ákos Kereszturi; Szaniszló Bérczi; András Sik; Tamás Pócs; Tibor Gánti; Eörs Szathmáry (February 2009). "Analysis of Dark Albedo Features on a Southern Polar Dune Field of Mars". Astrobiology. 9 (1): 90–103. Bibcode: 2009AsBio...9...90H (https://ui.adsabs.harvard.edu/abs/2009AsBio...9...90H). doi:10.1089/ast.2007.0212 (https://doi.org/10.1089%2Fast.2007.0212). PMID 19203240 (https://pubmed.ncbi.nlm.nih.gov/19203240).
- 41. F. Schmidt, S. Dout'e, B. Schmitt, Y. Langevin, J.P. Bibring and the OMEGA Team (2009). "Slab ice in the seasonal south polar cap of Mars" (http://meetingorganizer.copernicus.org/EPSC2009/EPSC2 009-521-2.pdf) (PDF). European Planetary Science Congress (EPSC) Abstracts. 4 (EPSC2009): 521–522. Retrieved 2 September 2009.
- 42. Möhlmann, Diedrich; Akos Kereszturi (5 January 2010). "Viscous liquid film flow on dune slopes of Mars". *Icarus*. **207** (2): 654. <u>Bibcode</u>:2010lcar..207..654M (https://ui.adsabs.harvard.edu/abs/2010lcar..207..654M). doi:10.1016/j.icarus.2010.01.002 (https://doi.org/10.1016%2Fj.icarus.2010.01.002).
- 43. "Martian spots warrant a close look" (http://www.esa.int/Science\_Exploration/Space\_Science/Exploring\_space/Martian\_spots\_warrant\_a\_close\_look). European Space Agency. 13 March 2002. Retrieved 8 September 2009.
- 44. Pócs, T.; A. Horváth; T. Gánti; Sz. Bérczi; E. Szathmáry (2003). ESA SP-545 Possible crypto-biotic-crust on Mars? (https://web.archive.org/web/20110721104940/http://www.colbud.hu/esa/publications/28CBC8ESASP-545pp265-266.pdf) (PDF). European Space Agency. Archived from the original (htt p://www.colbud.hu/esa/publications/28CBC8ESASP-545pp265-266.pdf) (PDF) on 21 July 2011. Retrieved 24 November 2008.
- 45. Gánti, Tibor; András Horváth; Szaniszló Bérczi; Albert Gesztesi; Eörs Szathmáry (31 October 2003). "Dark Dune Spots: Possible Biomarkers on Mars?". *Origins of Life and Evolution of Biospheres.* 33 (s 4–5): 515–557. Bibcode:2003OLEB...33..515G (https://ui.adsabs.harvard.edu/abs/2003OLEB...3 3..515G). doi:10.1023/A:1025705828948 (https://doi.org/10.1023%2FA%3A1025705828948). PMID 14604189 (https://pubmed.ncbi.nlm.nih.gov/14604189). S2CID 23727267 (https://api.semantic scholar.org/CorpusID:23727267).
- 46. Pócs, T.; A. Horváth; T. Gánti; S. Bérczi; E. Szathmáry (27–29 October 2003). <u>"38th Vernadsky-Brown Microsymposium on Comparative Planetology Are the dark dune spots remnants of the crypto-biotic-crust of Mars?" (https://web.archive.org/web/20110721104952/http://www.colbud.hu/esa/publications/26MosCBC10color.pdf) (PDF). Moscow, Russia. Archived from the original (http://www.colbud.hu/esa/publications/26MosCBC10color.pdf) (PDF) on 21 July 2011. Retrieved 7 September 2009.</u>
- 47. András Sik; Ákos Kereszturi. "Dark Dune Spots Could it be that it's alive?" (http://www.monochrom. at/dark-dune-spots/). Monochrom. Retrieved 4 September 2009. (Audio interview, MP3 6 min.)
- 48. Möhlmann, Diedrich T.F. (13 November 2009). "Temporary liquid water in upper snow/ice subsurfaces on Mars?". *Icarus*. **207** (1): 140–148. Bibcode:2010lcar..207..140M (https://ui.adsabs.harvard.edu/abs/2010lcar..207..140M). doi:10.1016/j.icarus.2009.11.013 (https://doi.org/10.1016%2Fj.icarus.2009.11.013).
- 49. Landis, Geoffrey A.; Oleson, Steven J.; McGuire, Melissa (9 January 2012). "Design Study for a Mars Geyser Hopper" (https://ntrs.nasa.gov/search.jsp?R=20120004036). Glenn Research Center. NASA. hdl:2060/20120004036 (https://hdl.handle.net/2060%2F20120004036). Retrieved 1 July 2012.
- Landis, Geoffrey A.; Oleson, Steven J.; McGuire, Melissa (9 January 2012), "Design Study for a Mars Geyser Hopper", <u>50th AIAA Aerospace Sciences Conference</u> (https://ntrs.nasa.gov/archive/nas a/casi.ntrs.nasa.gov/20120004036\_2012004260.pdf) (PDF), Glenn Research Center, NASA, AIAA-2012-0631, retrieved 1 July 2012

## **External links**

- Martian "Spiders" (http://www.martianspiders.com) photo repository.
- Arthur C. Clarke on "Martian Spider" features: 1 (http://www.martianspiders.com/Arthur%20C\_%20Cl arke%20Stands%20By%20His%20Belief%20in%20Life%20on%20Mars.htm)

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